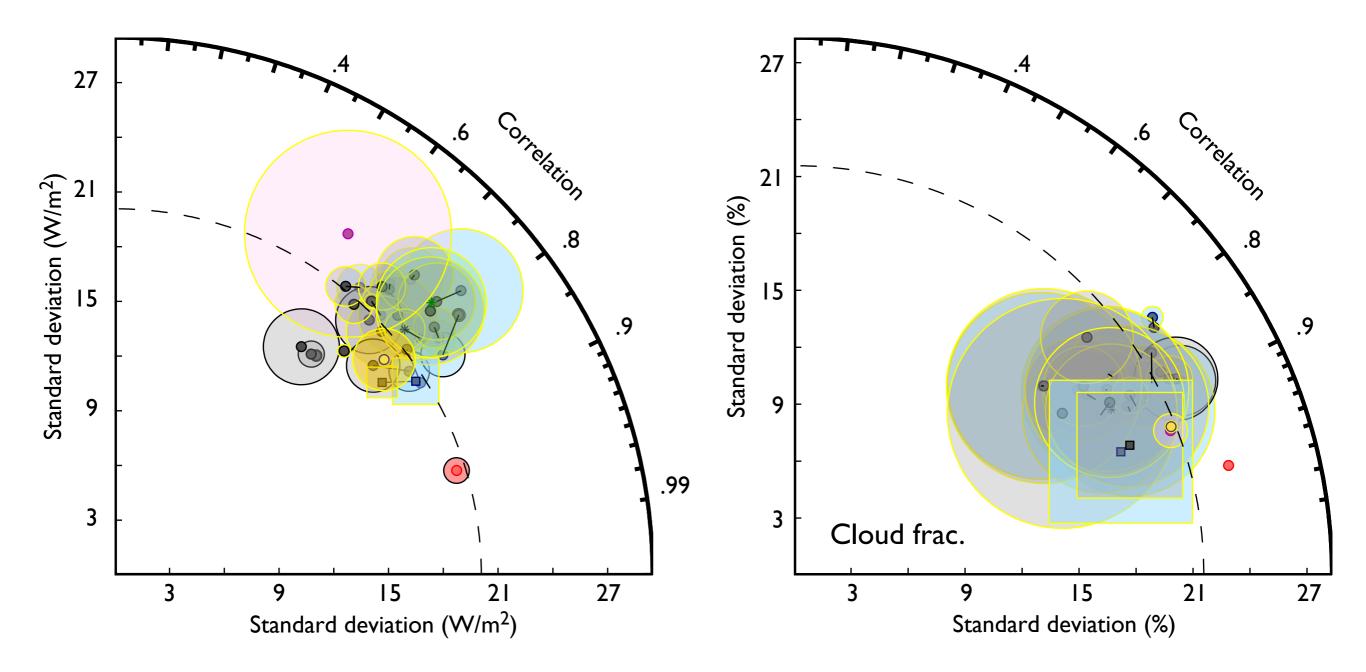
What we mean by "cloud:" The perils of evaluating climate models with satellite observations

Robert Pincus, University of Colorado

Steven Platnick, NASA/GSFC Steve Ackerman, University of Wisconsin Richard Hemler, NOAA/GFDL R. J. Patrick Hofmann, University of Colorado

Based on Pincus et al, manuscript submitted to J. Climate



Pincus et al., 2008

Making fair comparisons

We want to be able to attribute *differences* between observations and models to *model errors*

Some comparisons (e.g. temperature) can be direct, but satellites don't observe model cloud states

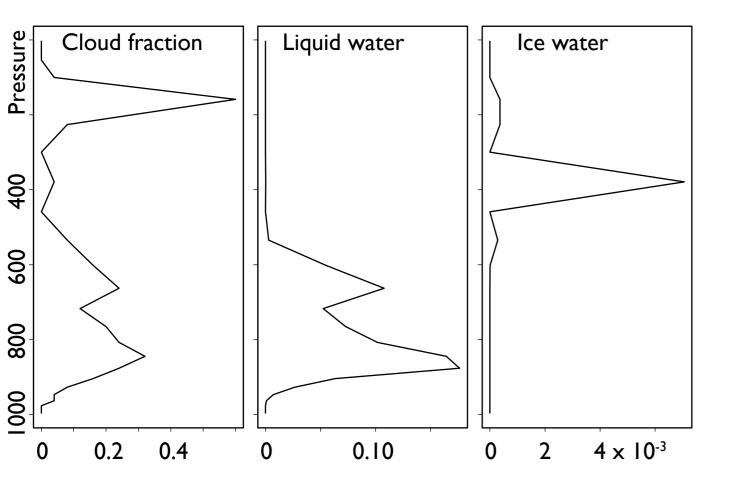
We require models of

sub-grid scale distribution of cloudiness

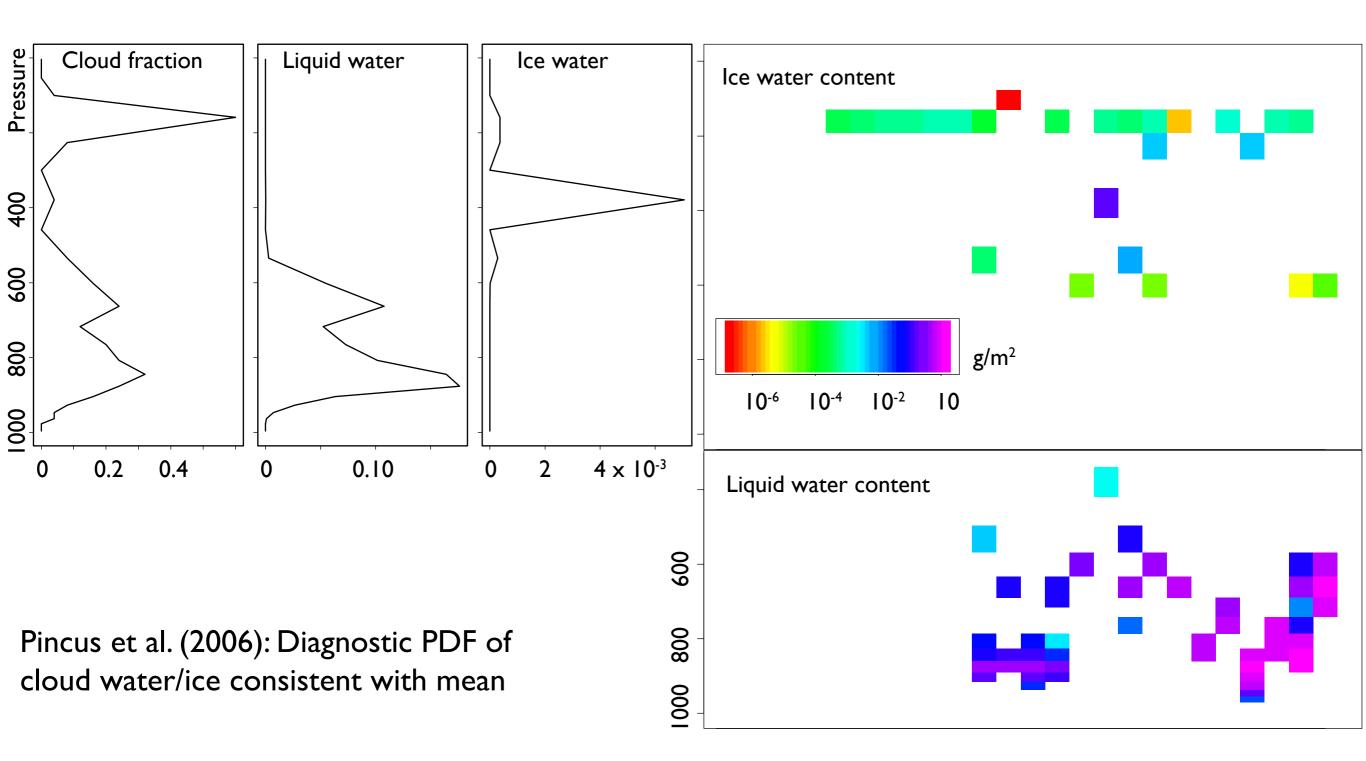
observational process

These are embodied in "instrument simulators" that operate on subcolumns

Sampling model state to treat fractional cloudiness/masking

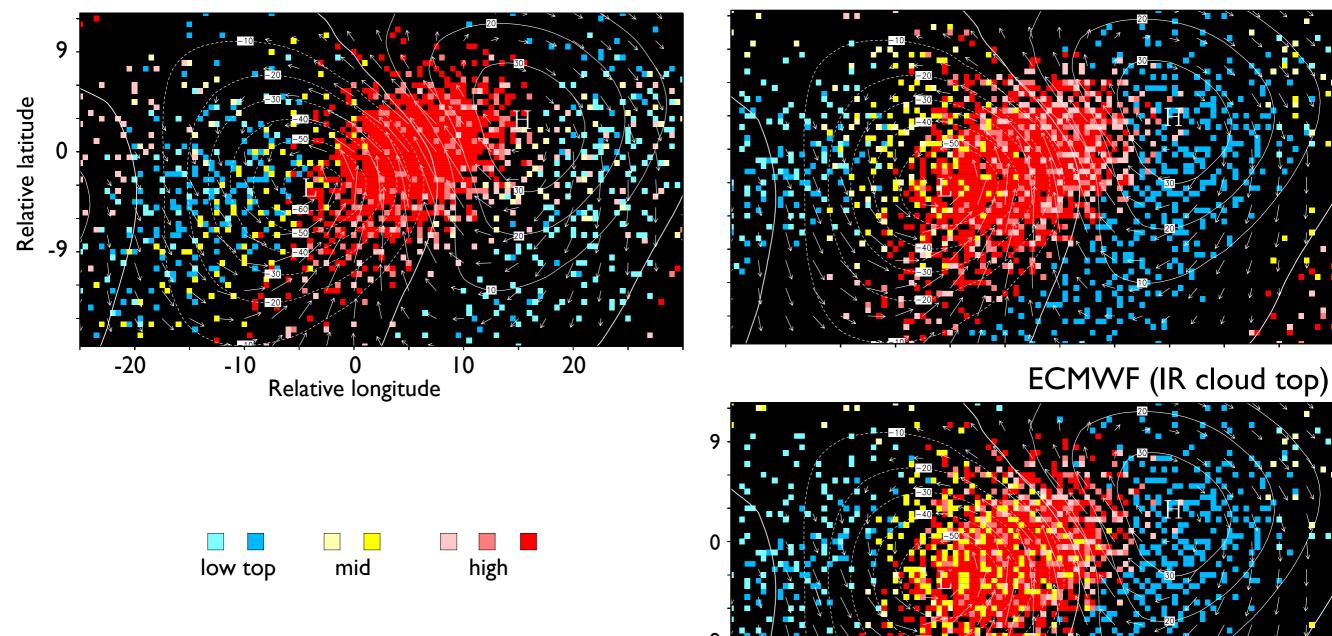


Sampling model state to treat fractional cloudiness/masking

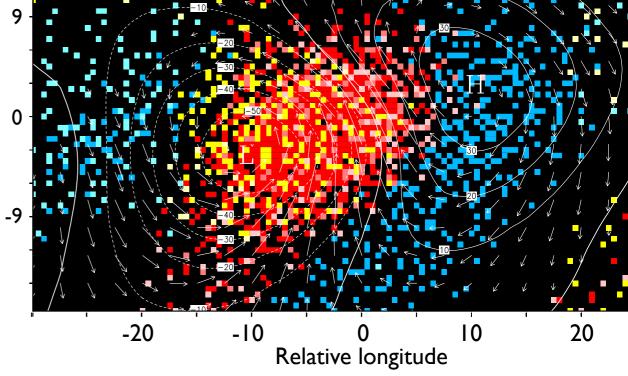








After Klein and Jakob, MWR, 1999



A MODIS simulator for climate models (i)

Accepts sub-column inputs of $r_{e(l,i)}(z)$, $\tau_{(l,i)}(z)$ or $q_{(l,i)}(z)$

Provides subcolumn estimates of

$$p_c = \int_{TOA}^{\tau=1} p(z)\sigma_c(z)dz$$
 (when > 700 mb, use ISCCP IR)

$$P = \int_{\text{TOA}}^{\tau=1} P(z)\sigma_c(z)dz \qquad \text{(can be "undetermined")}$$

 $au = \int_{TOA}^{sfc} \sigma_c(z) dz$ (no errors, as ISCCP simulator)

 $r_e = F^{-1}(F(r_e(z)))$ (pseudo-retrieval based on near-IR fluxes)

A MODIS simulator for climate models (ii)

Estimate liquid, ice water path from optical thickness, particle size

Aggregate sub-columns

cloud fractions (total, liquid, ice, high, middle, low) in-cloud linear means of all quantities in-cloud logarithmic mean for optical thickness optical thickness/cloud top pressure joint histogram (18 statistics total) A MODIS simulator for climate models (iii)

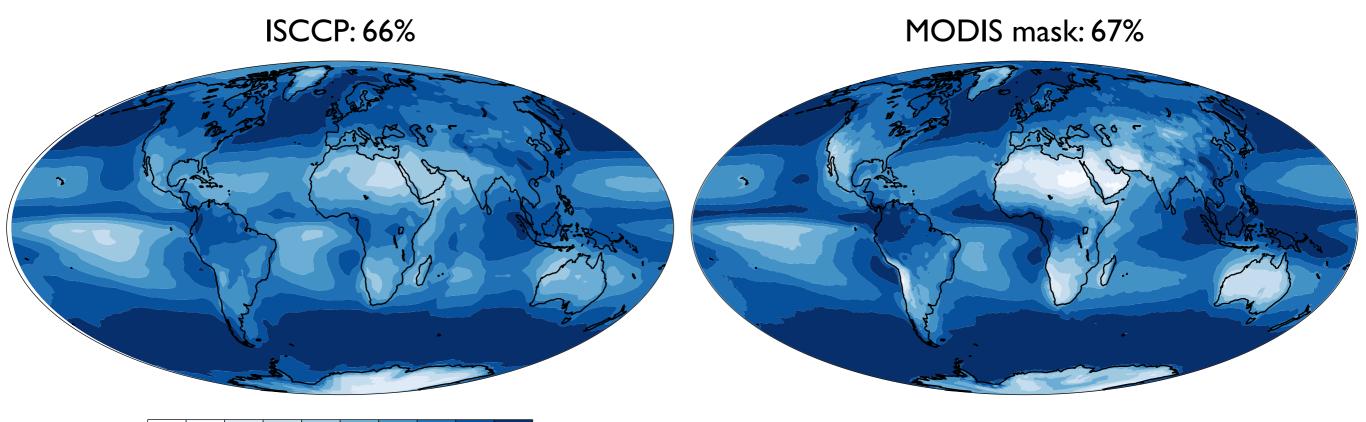
MODIS simulator distributed as part of the larger COSP package

Available in essentially every GCM, though MODIS output is not required for CMIP5

We provide a customized data set for comparisons

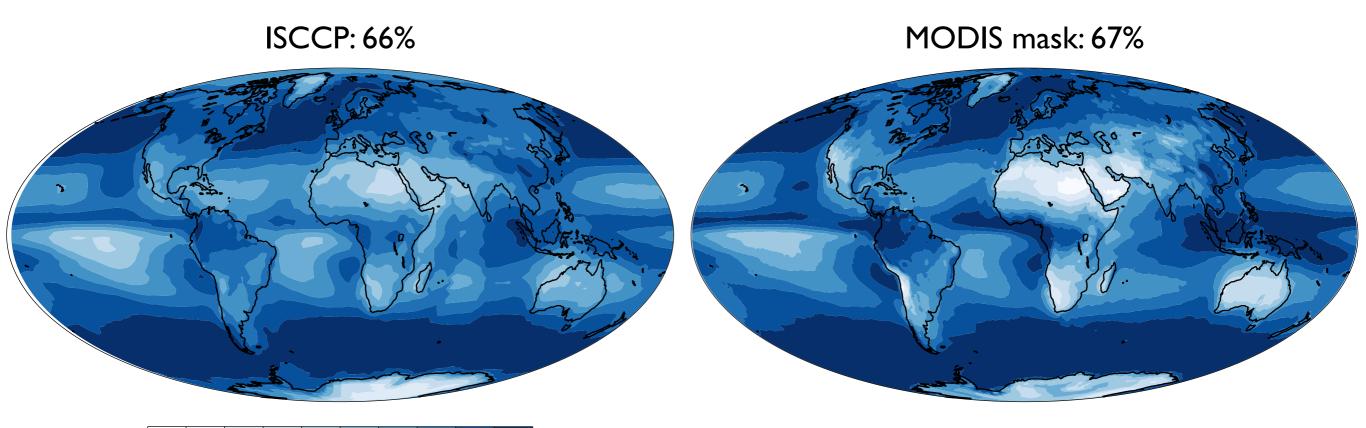
Included results from cloud mask and cloud retrievals - and that's where the fun begins

How much of the planet is cloudy?



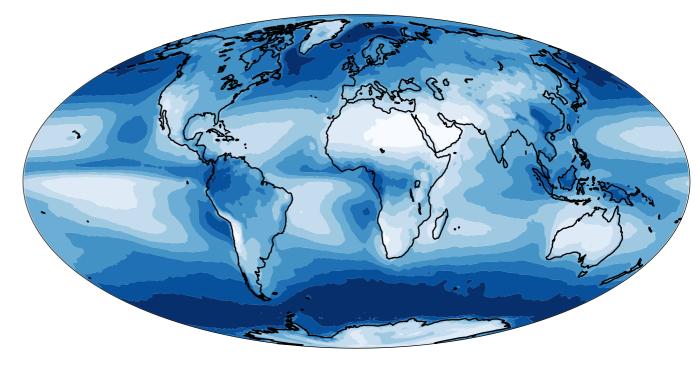
I	0	Clc	bud	fra	cti	on	(%)	8	0

How much of the planet is cloudy?



0	Clc	bud	fra	cti	on	(%)	8	0

MODIS retrievals: 50%

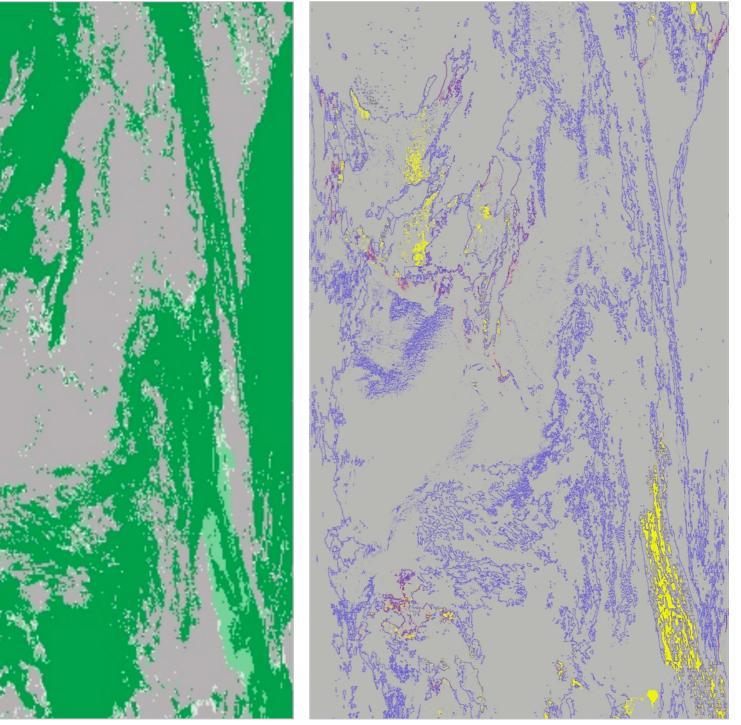


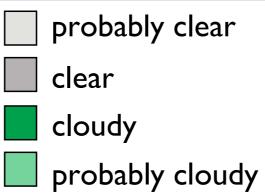
SWIR composite



Cloud Mask overall conf.

"Clear Sky Restoral"







spatial/spectral tests edge detection

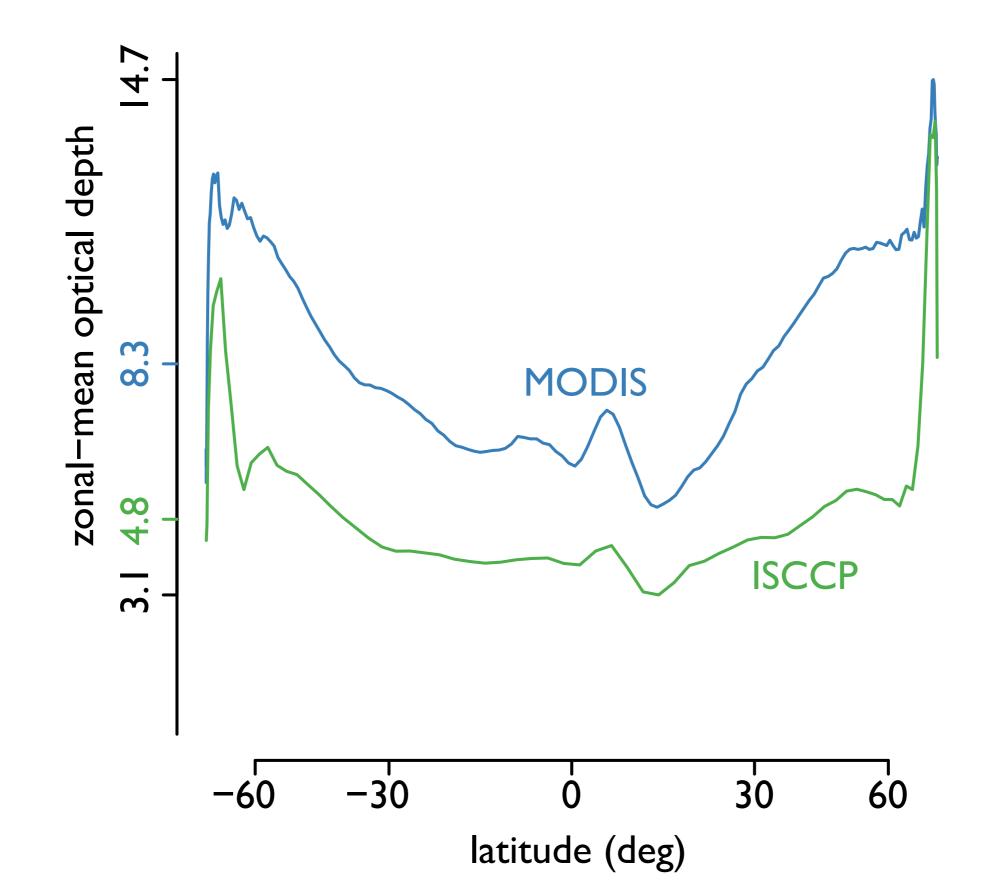


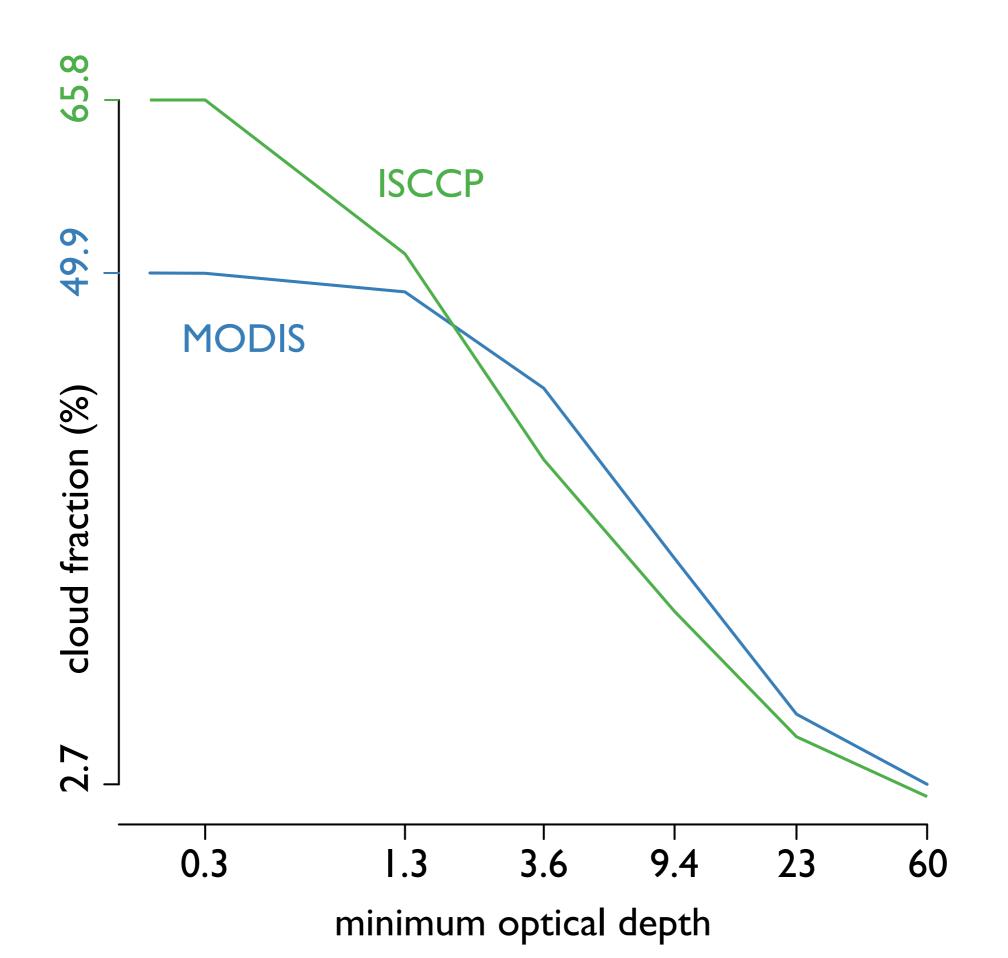
250m cloud mask

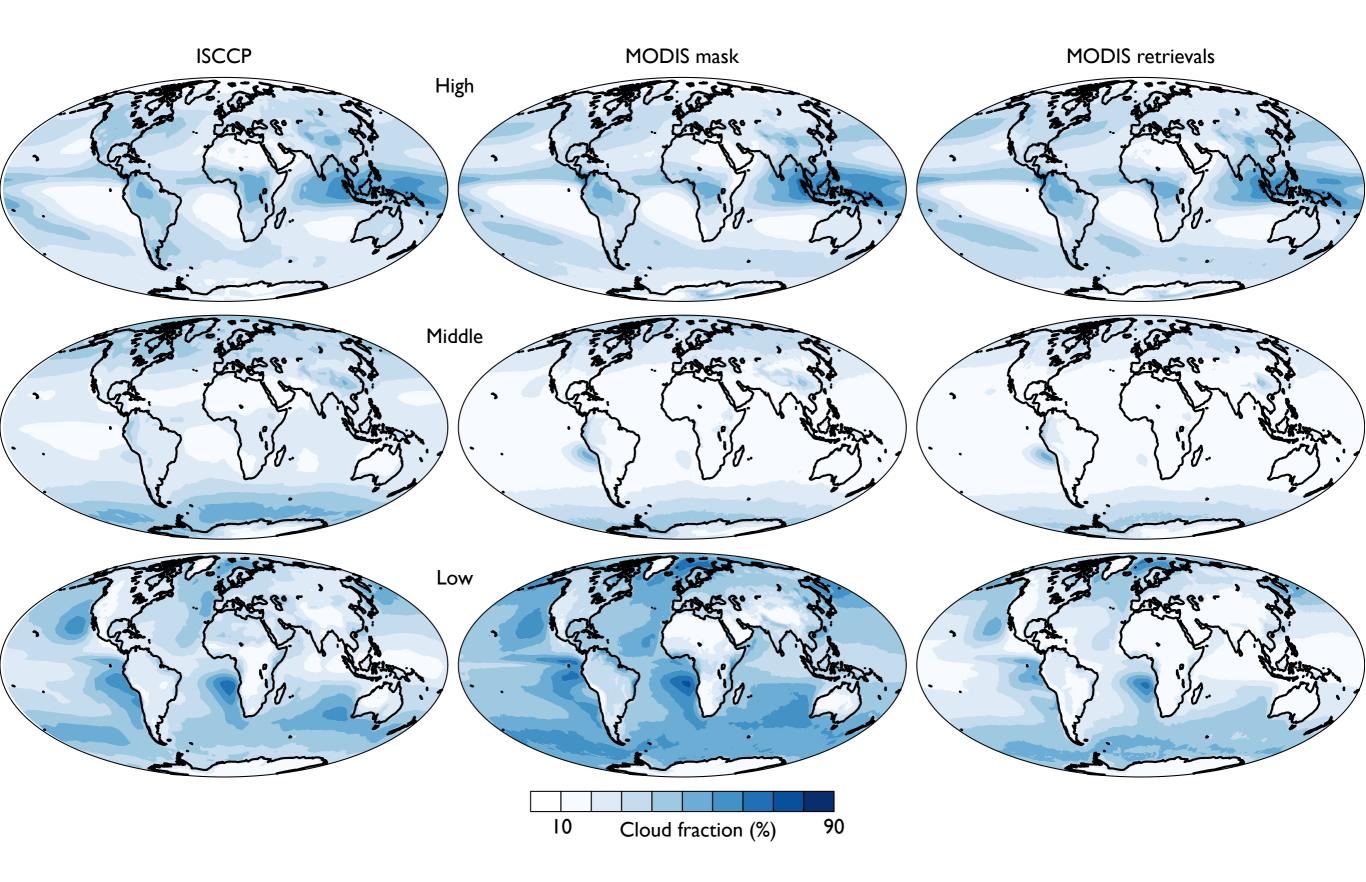
Clear-sky restoral was introduced in Collection 5

One practical goal was to reduce the very large particle sizes obtained from MODIS 2.1 μ m observations, but...

Clear-sky restoral didn't change average particle size values much because retrievals at cloud edges were failing







Observation

Pixels are removed by clear-sky restoral mostly because they are near cloud edges or are inhomogeneous at 250 m scale

This population is

nearly all the clouds observed by ISCCP with au < 1.3

assigned high cloud top pressure by MODIS but distributed through the atmosphere by ISCCP (~1/3 are consistent with failed retrievals by ISCCP)

Interpretation

Roughly 15% of the planet is covered by clouds less than I km in size

Omitting these pixels leads to truncation errors

Literal interpretations of retrievals is inappropriate

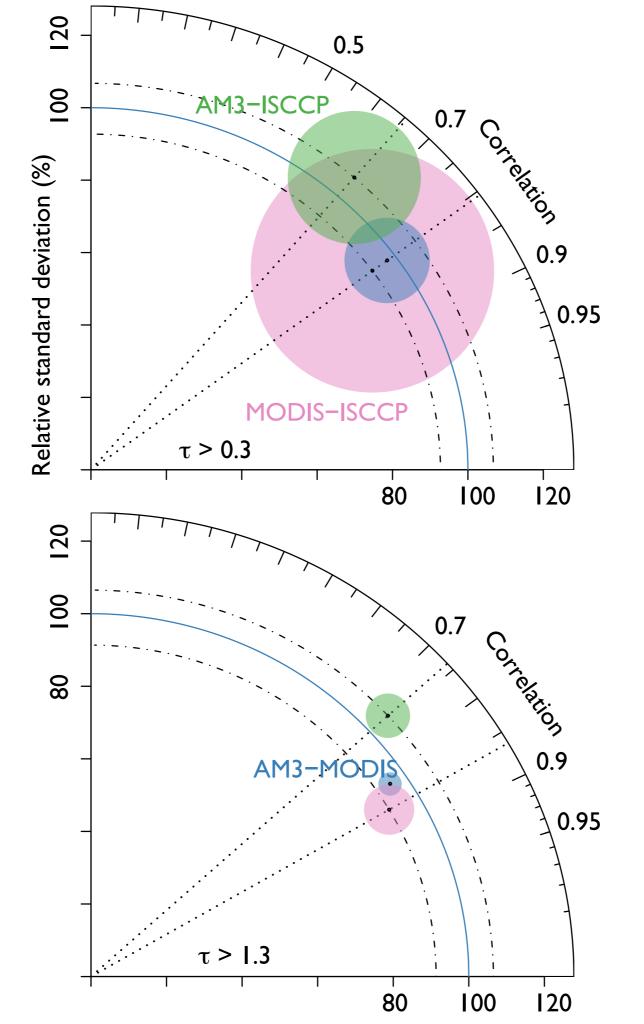
Implication

Cloud fraction estimates explicitly depend on detector resolution and sensitivity

Large-scale models have no concept of spatial scale below the grid size

Total cloudiness is a fragile basis for comparison

Comparisons among observations (and between models and observations) are fair only when the same population is included



Working with the climate modeling community means

doing some translation

understanding our observations more fully

acknowledging when we simply can't speak the same language