




SNPP VIIRS Lunar Images and Irradiances — a Correlation Study of Moon Image Orientation

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VIIRS Observations of the Moon

SNPP and NOAA-20 execute roll maneuvers each month to capture the Moon in the Earth-view sector at phase angle $\sim 51^\circ$ before Full Moon

- when the Moon is observable, typically October through June
- as the spacecraft traverses its orbit, the Moon passes through the field of view 
- roll angle specified to center the Moon disk in Earth view:



SNPP VIIRS image d20170604_t1934579, band M7, scan 12

- centering avoids stray light, seen by stretching the display level:



Lunar Calibration

USGS lunar calibration works with lunar irradiances, comparing sensor measurements against reference values generated by the ROLO model.

- Reference irradiances are computed for the Sun-Moon-observer geometry (phase, librations, distances) corresponding to the instrument's Moon observations, transformed to the sensor's band wavelengths
- Irradiance measurements from Moon images involves spatial integration of pixels on the lunar disk:

$$E_{\text{meas}} = \frac{\Omega_p}{\eta} \sum_i^N L_i$$

Ω_p = pixel IFOV (solid angle)

η = oversampling factor

L_i = pixel radiance

N = # of pixels on Moon

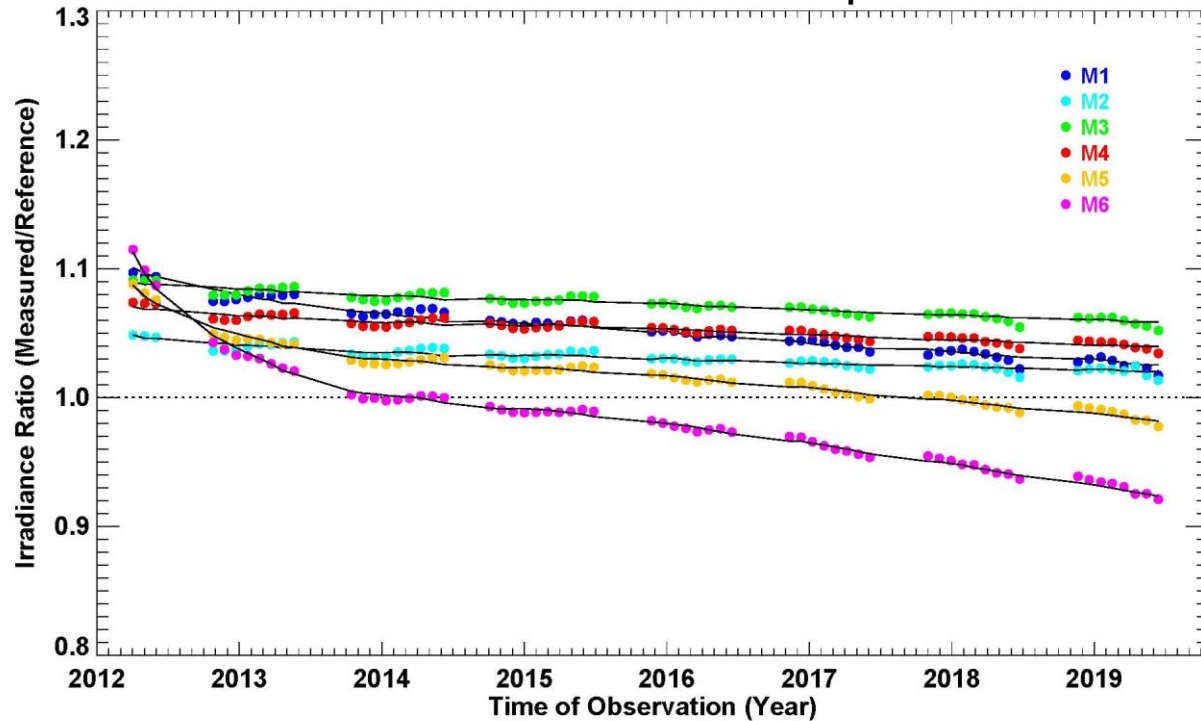
- ★ The accuracy of lunar irradiance measurements from images depends on careful evaluation of:
 - net radiance: subtraction of the dark background
 - actual detector spatial response (IFOV, different from GSD)
 - oversampling of the Moon disk (different from slew/sampling)

Lunar Calibration

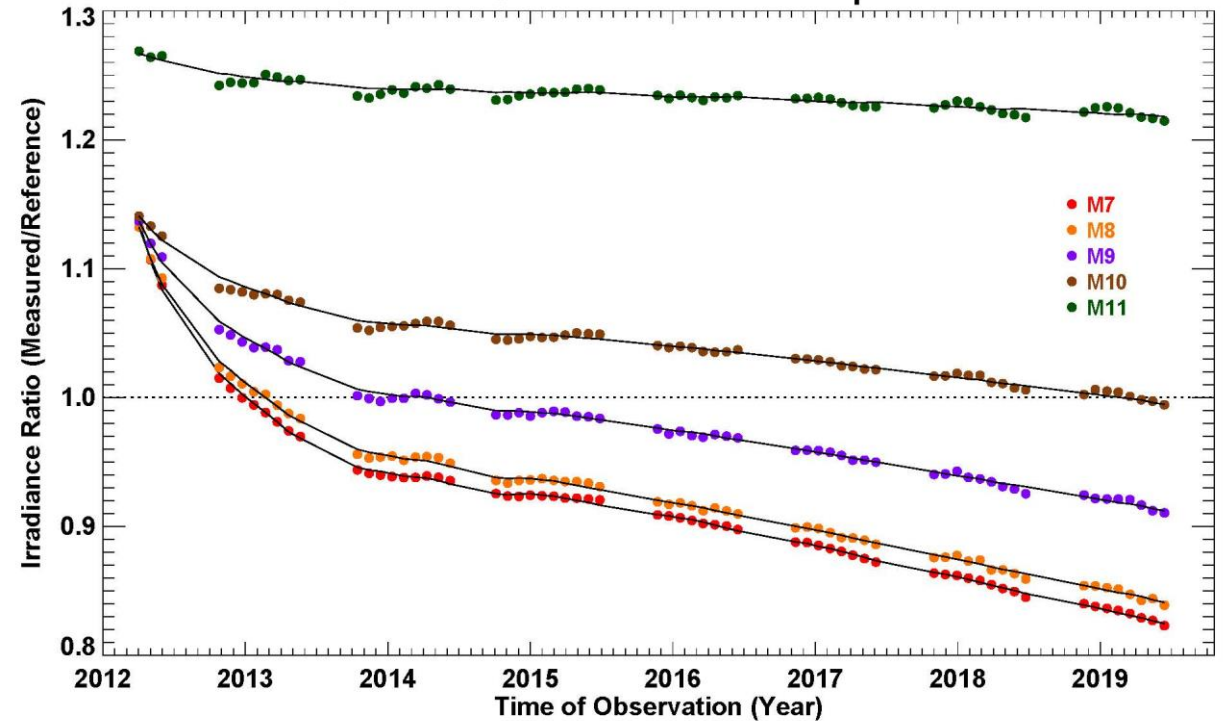
Typical usage: tracking sensor response changes on orbit

- time series of measurement/model ratios reveal sensor response trends

SNPP VIIRS Lunar Irradiance Comparisons



SNPP VIIRS Lunar Irradiance Comparisons



VIIRS Lunar Image Analysis at USGS

Moon image processing to irradiance, independent of VCST and OBPG

- RDRs obtained from NOAA CLASS
- SDR software system installed on USGS compute cluster: ADL 4.2.8
 - code modifications to remove time-dependent calibration components; all granules processed identically
- VIIRS moon image spatial integration routines developed at USGS
 - detector dark level evaluation from deep space regions around the Moon disk
 - radiance conversion using SDR base calibration factors, extracted from SDR files
 - pixel solid angle derived from Horizontal Sampling Interval (from geolocation ATBD)
- temporal response trends corrected using daily-average SD f-factors for days of Moon observations (thanks to VCST)

USGS results: time series of measurement/ROLO ratios, **not normalized**



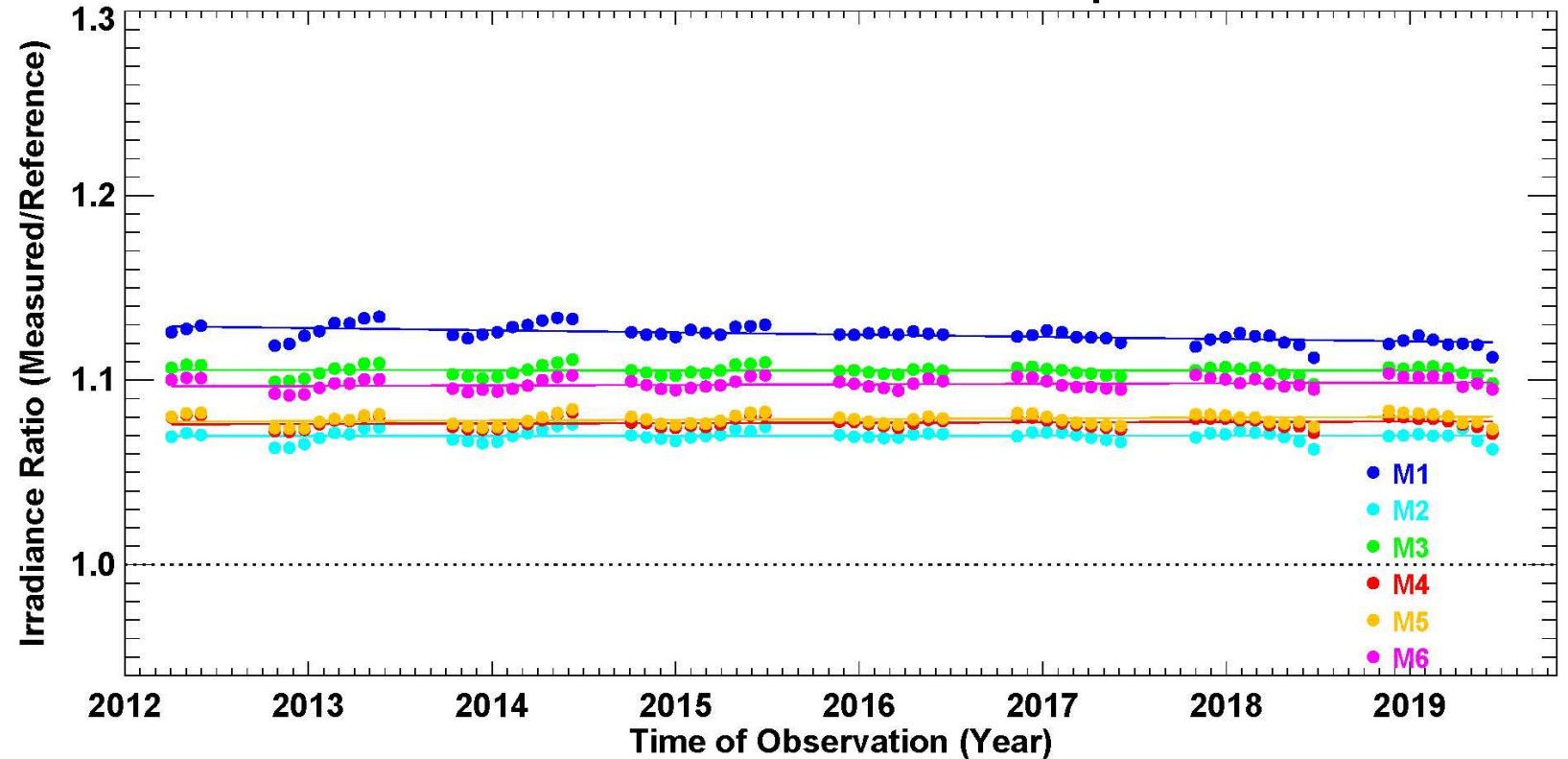
USGS Results for VIIRS Lunar Calibration

Lunar irradiance ratios (VIIRS/ROLO)
— de-trended —

Notable features:

- discrepancies (offsets) 7-16% for M1 to M6
- **band-correlated oscillation pattern**

SNPP VIIRS Lunar Irradiance Comparisons



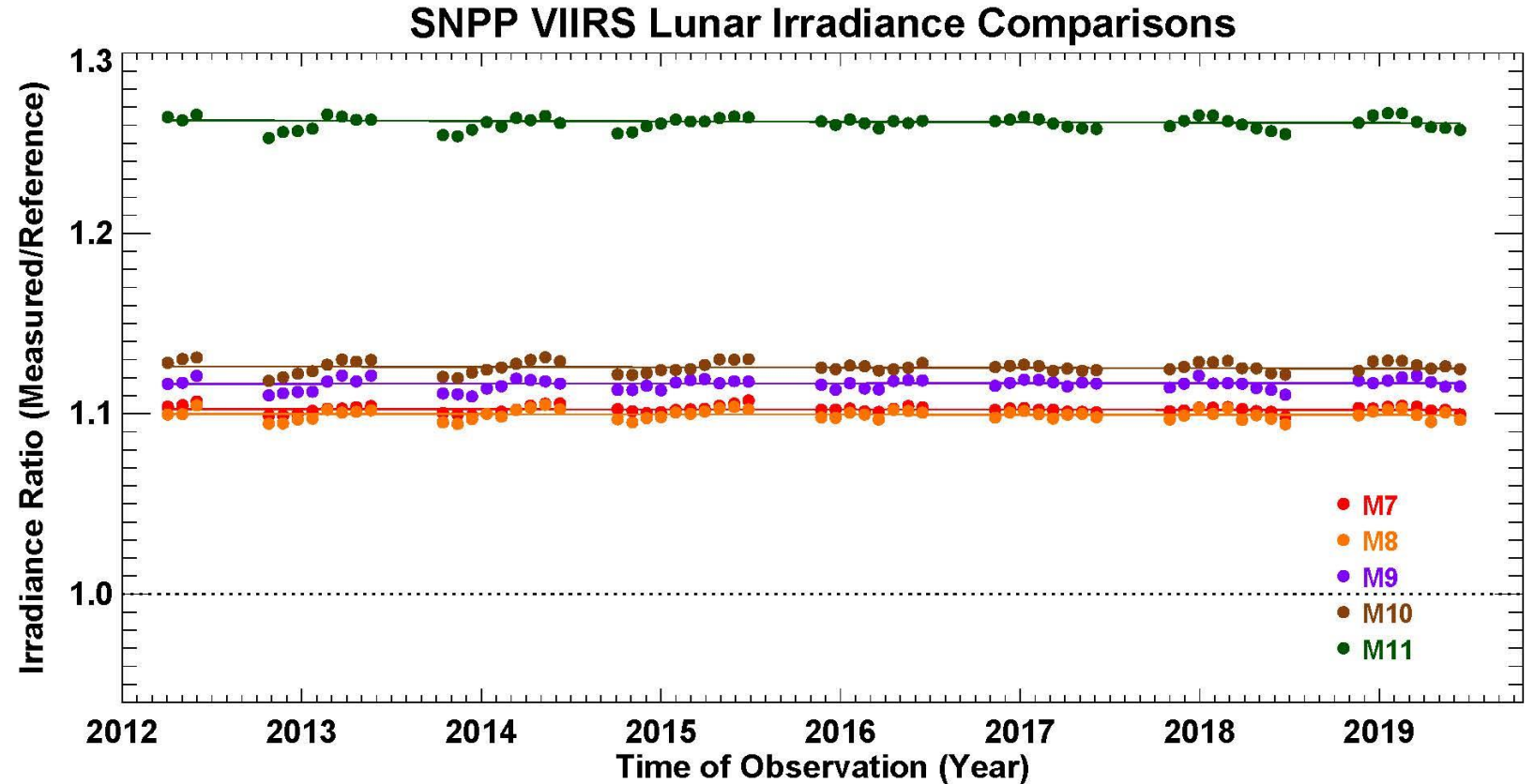
- lines show linear fits to time series
 - slopes reveal residual temporal drifts (small)

USGS Results for VIIRS Lunar Calibration

Lunar irradiance ratios (VIIRS/ROLO)
— de-trended —

Notable features:

- discrepancies (offsets)
9-14% for M7 to M10
~26% for M11
- **band-correlated oscillation pattern**



- lines show linear fits to time series
– slopes reveal residual temporal drifts (small)

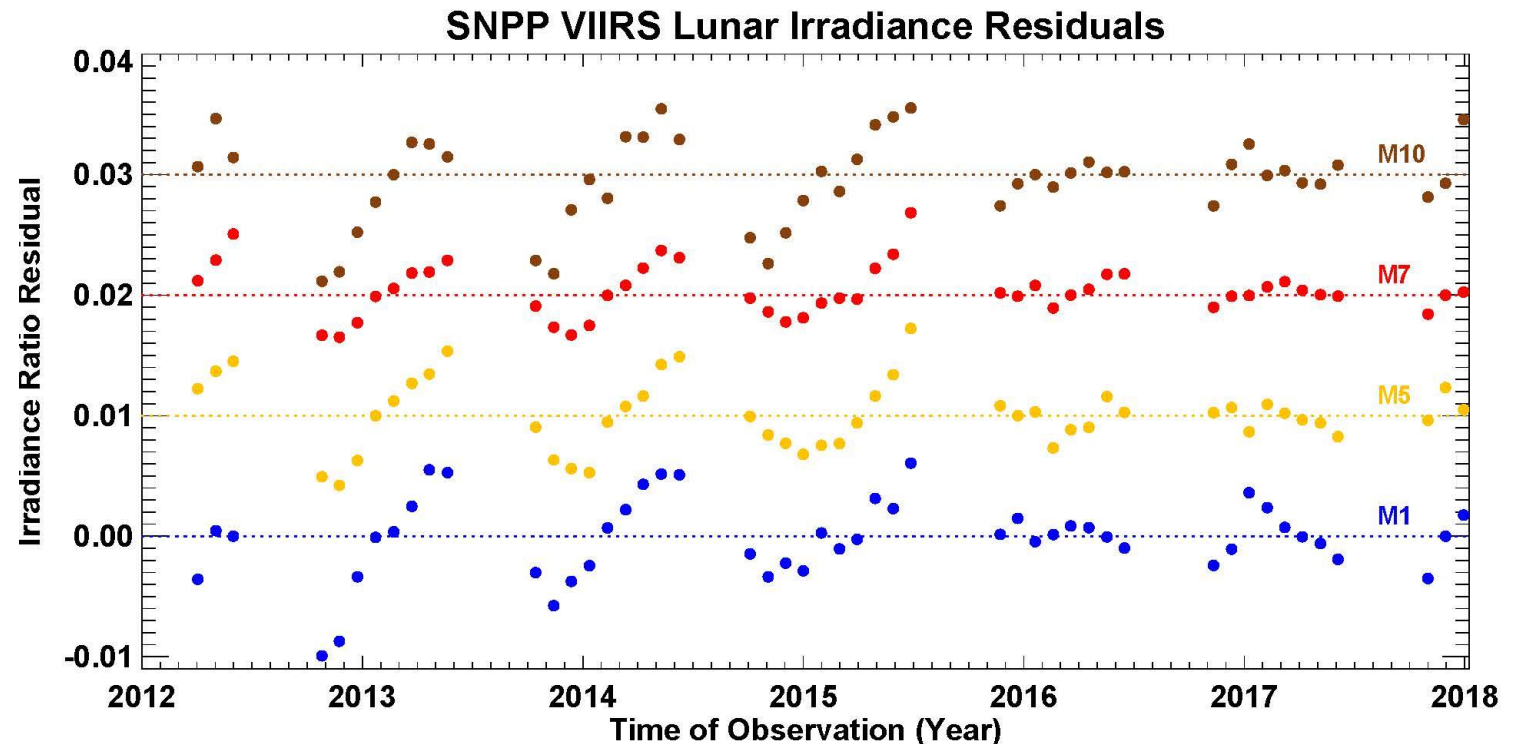
Investigating the Oscillation Patterns

The cross-band correlation suggests an origin in the ROLO lunar model.

- for a given sensor band, the model results are governed by only the phase angle and lunar librations
 - the only significant correlations were found with the observer (sensor) lunar librations

Analysis approach:

study residuals to the
linear trend fits ➔
(offset for clarity)



Investigating the Oscillation Patterns

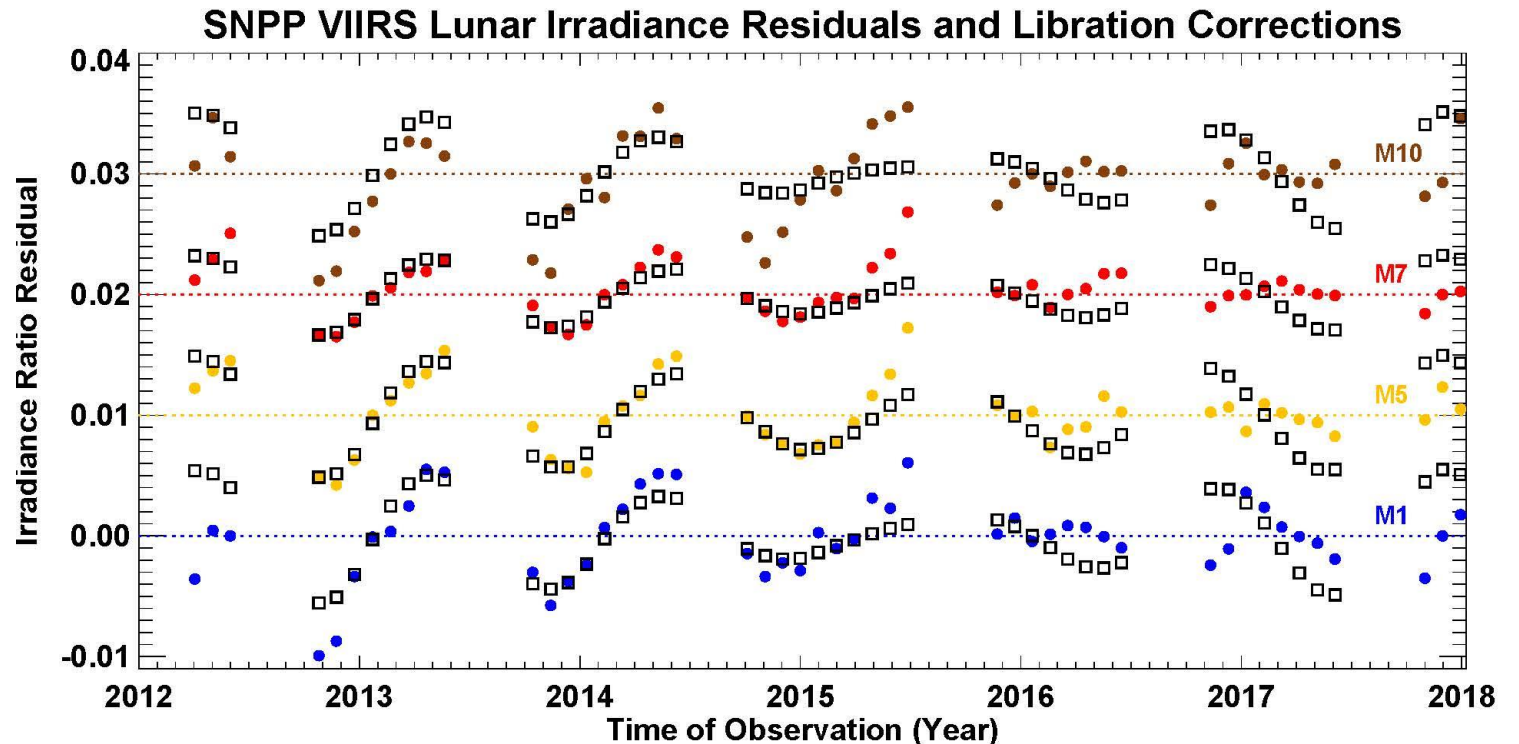
Developed correction to the ROLO model irradiances — a combined linear function of observer libration longitude and latitude:

$$E'_{\text{ref}} = E_{\text{ref}} (1 + c_0 \phi + c_1 \theta)$$

ϕ = sub-observer longitude

θ = sub-observer latitude

- corrections scaled to irradiance residuals →
- good correlation for years 2013-2015, then breaks down
- not a valid approach



★ *A correction to the ROLO model must specify a property of the Moon, and thus be valid for all observations by all instruments*

Investigating VIIRS Image Integration to Irradiance

VIIRS Moon images are presumed to be neither oversampled nor undersampled.

- if oversampling=1.0 does not strictly hold, then the irradiance measurements should show a dependence on the Moon image orientation in relation to the along-scan direction
 - due to the distribution of radiance across consecutive frames

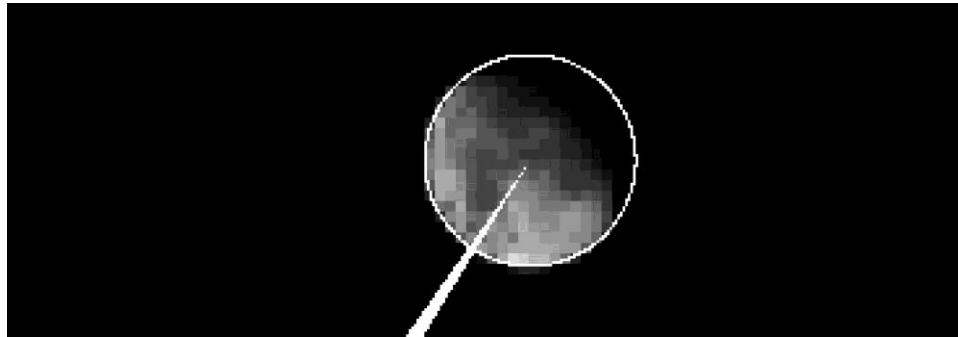
Image orientations were determined for VIIRS I-band images by selecting high-contrast pixels that define the lunar bright limb.

Investigating VIIRS Image Integration to Irradiance

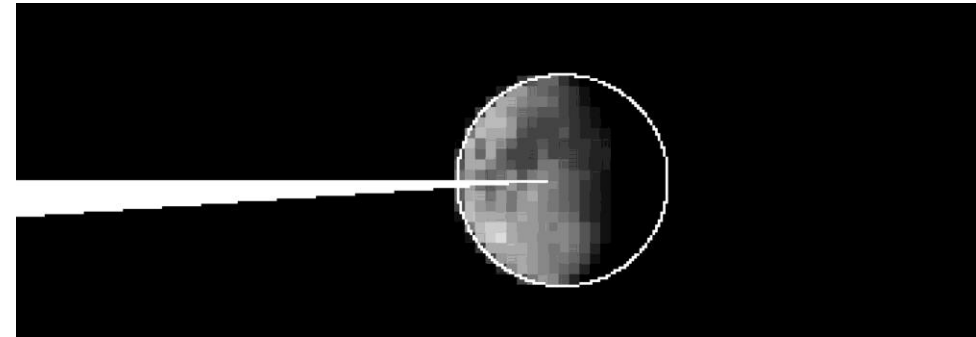
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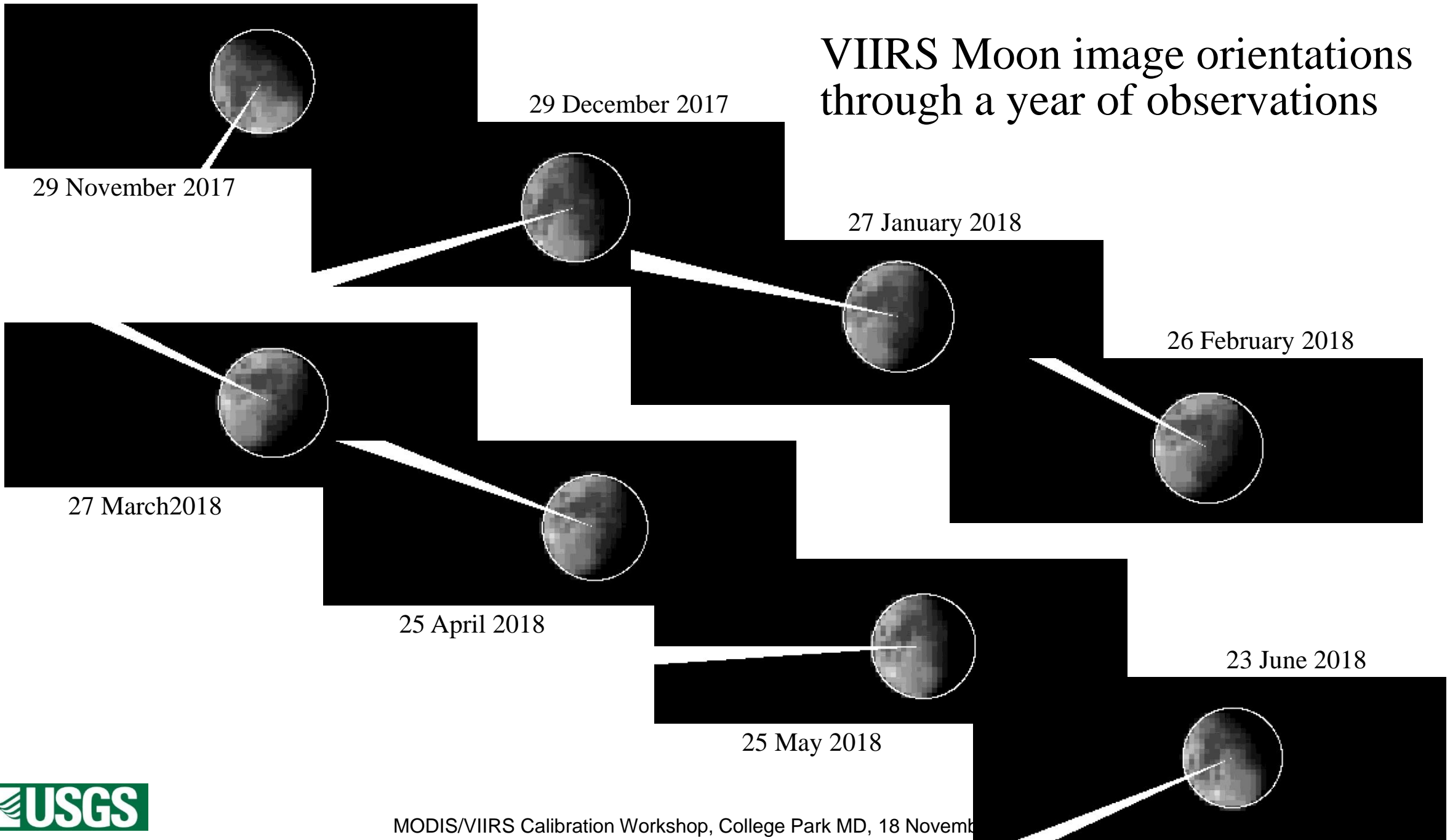


29 November 2017 236.6° ccw from +X



25 May 2018 181.9° ccw from +X

VIIRS Moon image orientations through a year of observations



Study Results and Conclusions

- Examining the complete series of VIIRS Moon observations shows the image orientations have an annual repeat cycle
 - the oscillations seen in the lunar measurement/model residuals have a different periodicity, therefore this effect cannot be attributed to image orientation
 - this negative result does not rule out potential along-scan oversampling errors
 - the cause of the oscillations has not conclusively been determined
- Other aspects of VIIRS lunar image spatial integration to examine:
 - motion of the Moon relative to the line of sight during scans
 - possible effects of pixel aggregation
- The oscillation effect is small, ~1% or less, but coupled across VIIRS bands
 - thus given consideration in the ROLO model redevelopment effort

Thank You!

